

# CROME experiment and its results

Jan Pękala  
Instytut Fizyki Jądrowej PAN



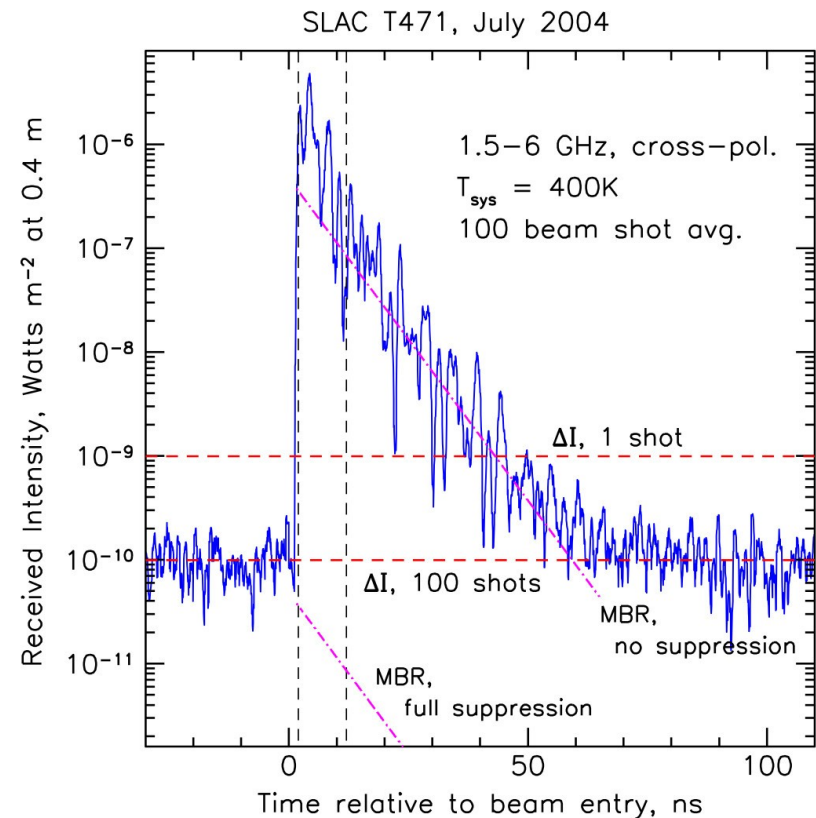
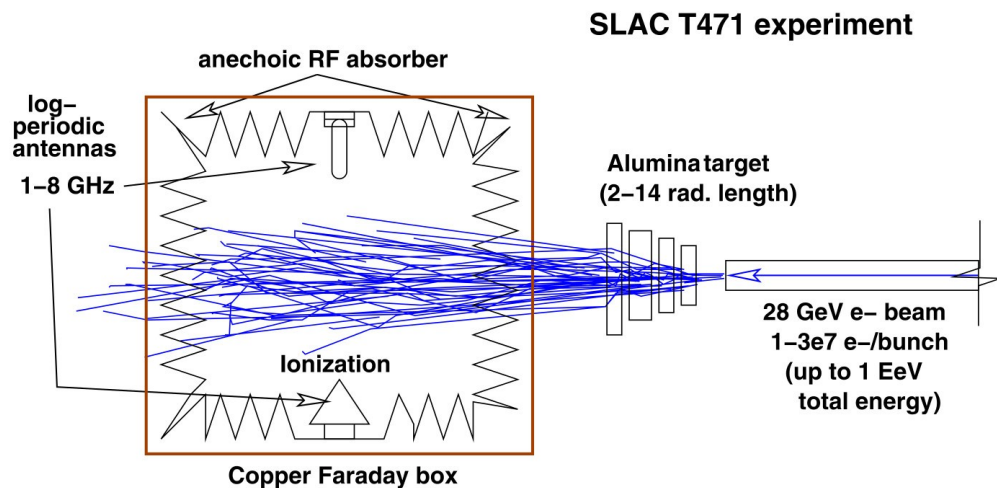
CROME – Cosmic-Ray Observation via Microwave Emission

# Molecular bremsstrahlung from air showers?

Experimental evidence for molecular bremsstrahlung radiation has been presented (Gorham et al., Phys Rev D 78 (2008) 032007).

The passage of charged particles of air shower produces plasma in the air. Free electrons could emit microwave photons.

⇒ **Isotropic, unpolarized microwave emission should be detectable.**



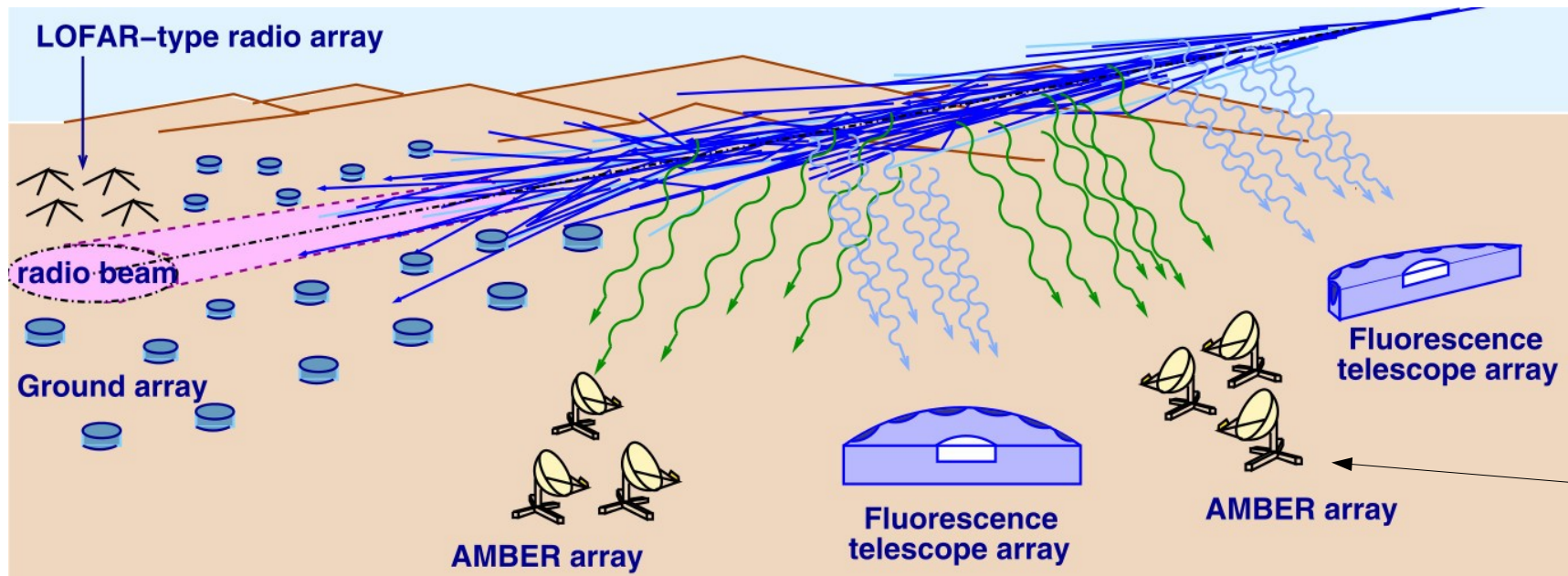


# New technique of air shower observation?

Possibility of developing a new method of air shower observations:

- very low natural background,
- transmission not obstructed by atmosphere,
- observation possible 24 hrs/day,
- easily adaptable technology available (satellite TV).

**Experimental tests needed for confirmation.**



# Cosmic-Ray Observation via Microwave Emission

## Location and external trigger

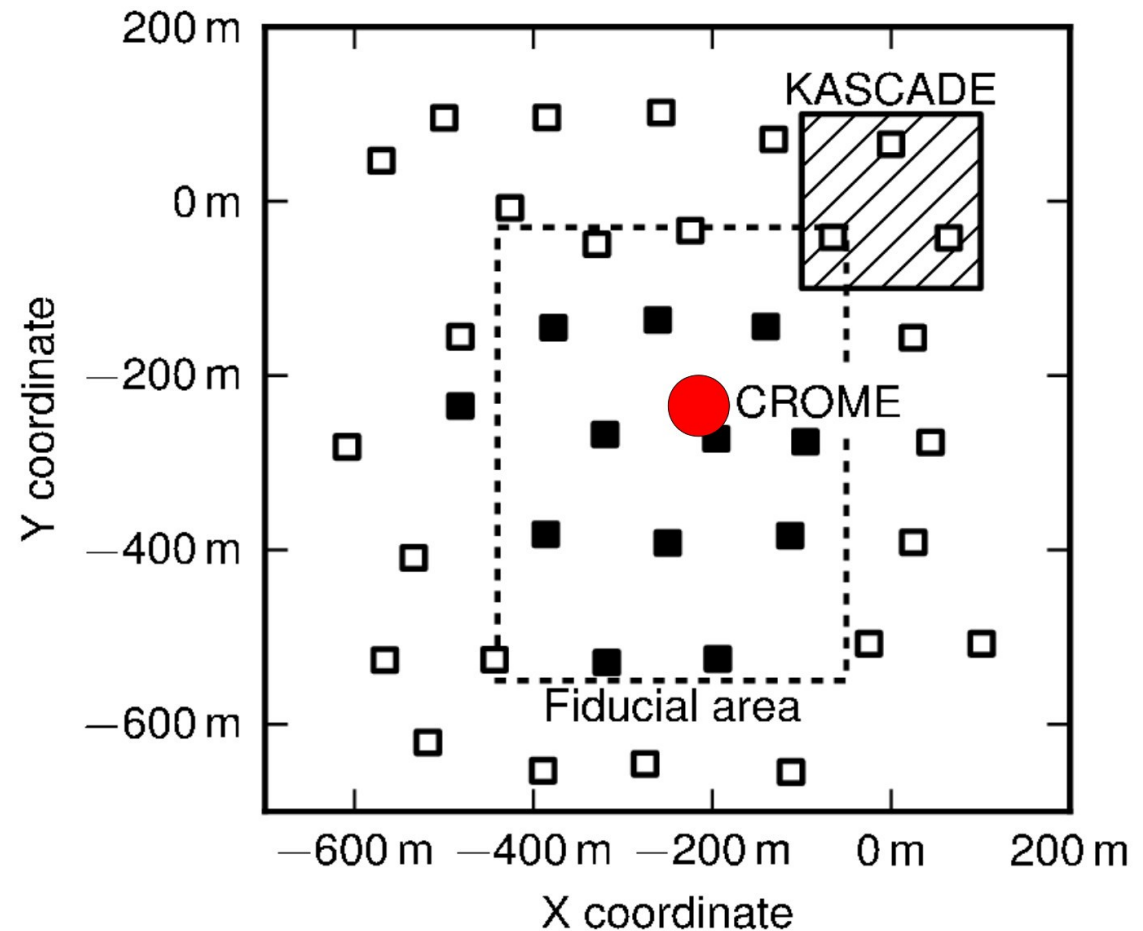
**CROME experiment was operating with KASCADE-Grande trigger:**  
May 2011 – November 2012

Array of radio antennas recording signal in several bands.

Triggered by air showers with energies  $10^{15} - 10^{18}$  eV.

Precise reconstruction of showers detected by KASCADE-Grande:

- $0.8^\circ$  for the arrival direction,
- 6 m for the position of shower axis,
- 20% for the energy.



CROME located in the center of KASCADE-Grande (Karlsruhe Institute of Technology).



# Overview of the CROME antennas

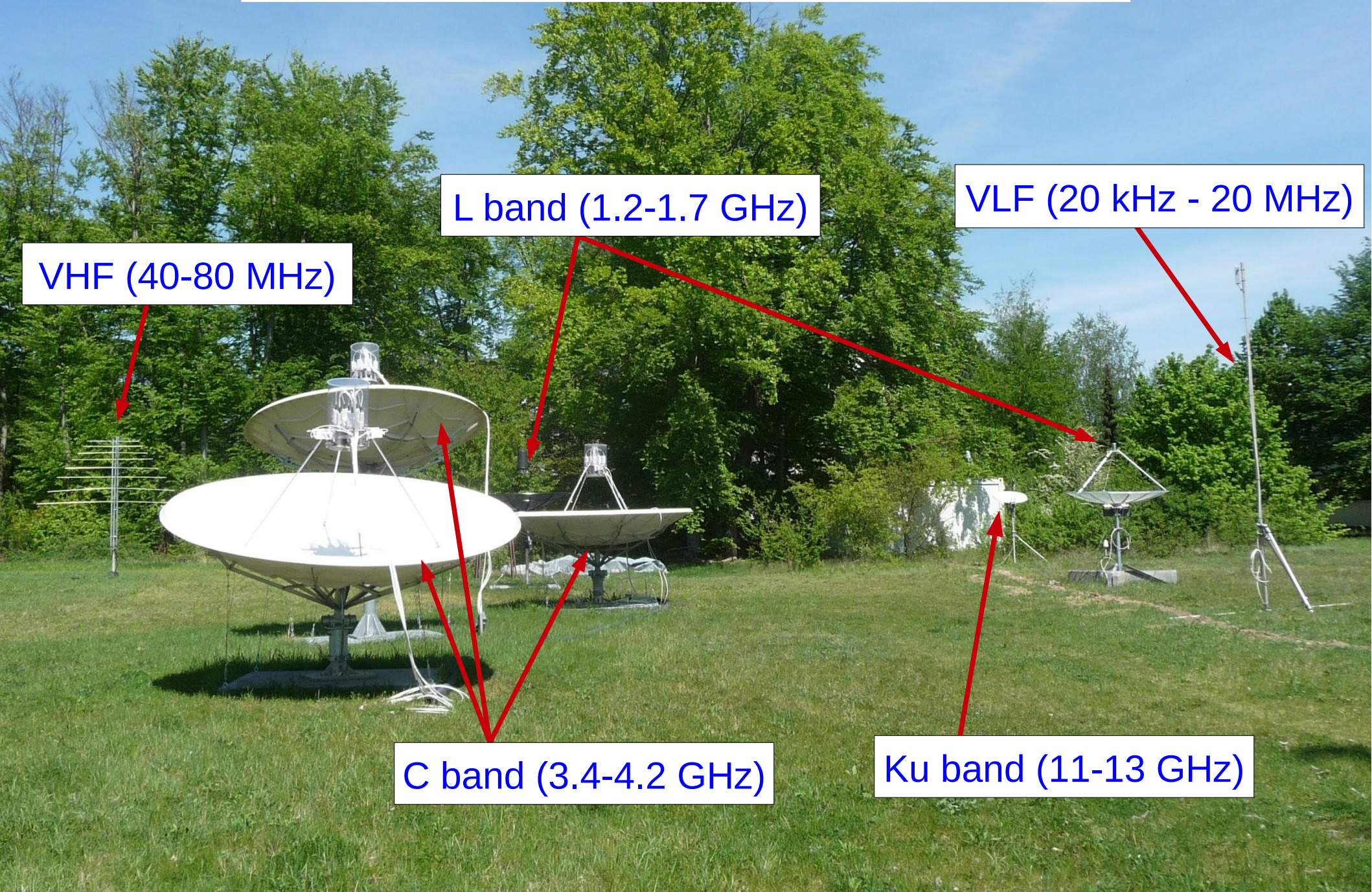
VHF (40-80 MHz)

L band (1.2-1.7 GHz)

VLF (20 kHz - 20 MHz)

C band (3.4-4.2 GHz)

Ku band (11-13 GHz)





# Observation in the C band (3.4 – 4.2 GHz)

The C band is used for satellite communication.  
Low natural background.  
Negligible atmospheric attenuation.

Commercial parabolic reflectors (335 cm diameter).

Antennas pointed upward – minimization of the distance to the shower maximum.

Two of the antennas tilted by  $15^\circ$  North and South to observe showers with different angles relative to the local geomagnetic field.

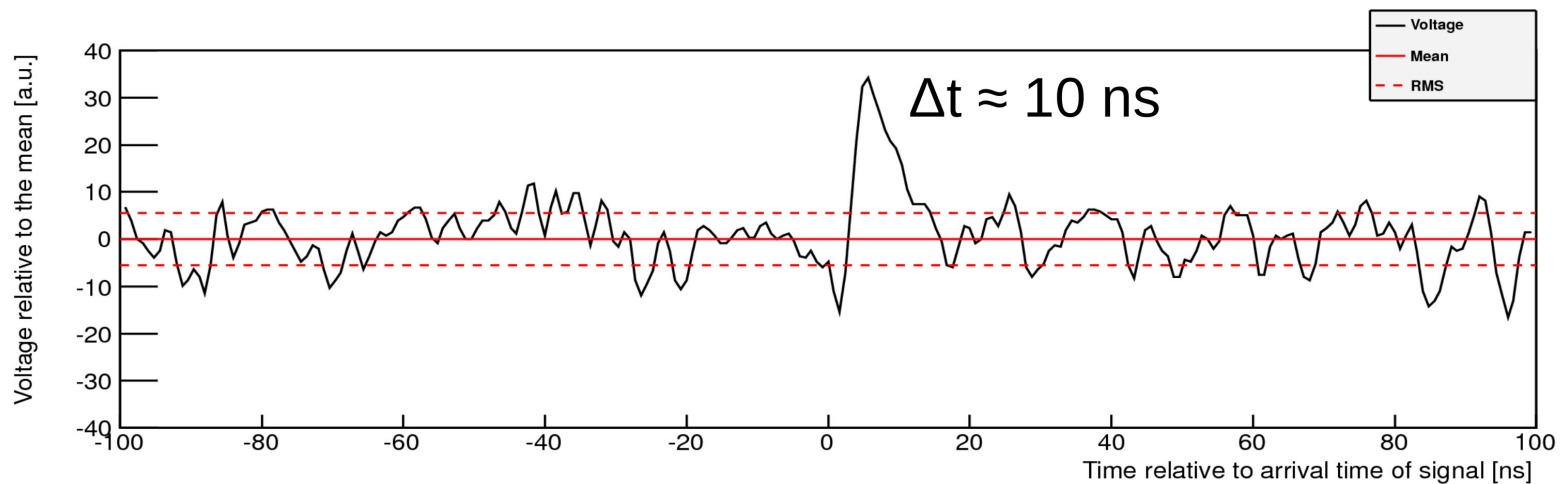
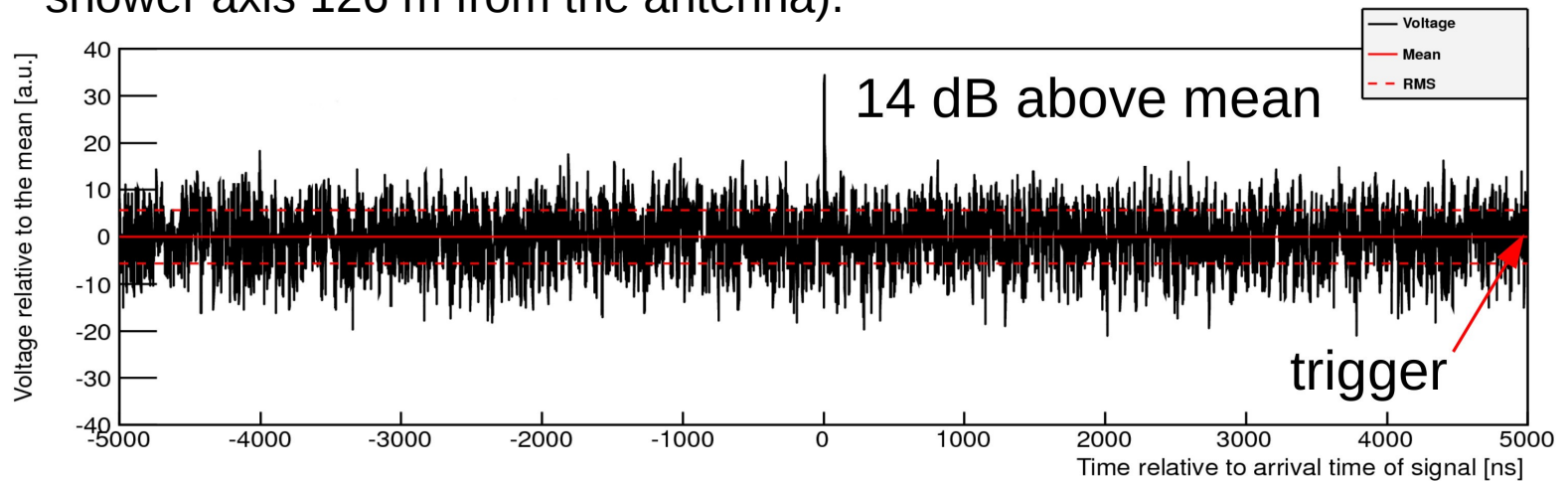
13 linearly polarized receivers in a camera.



# Finding microwave signal

Readouts were recorded for a time of  $10\ \mu\text{s}$  before and after each KASCADE-Grande trigger.

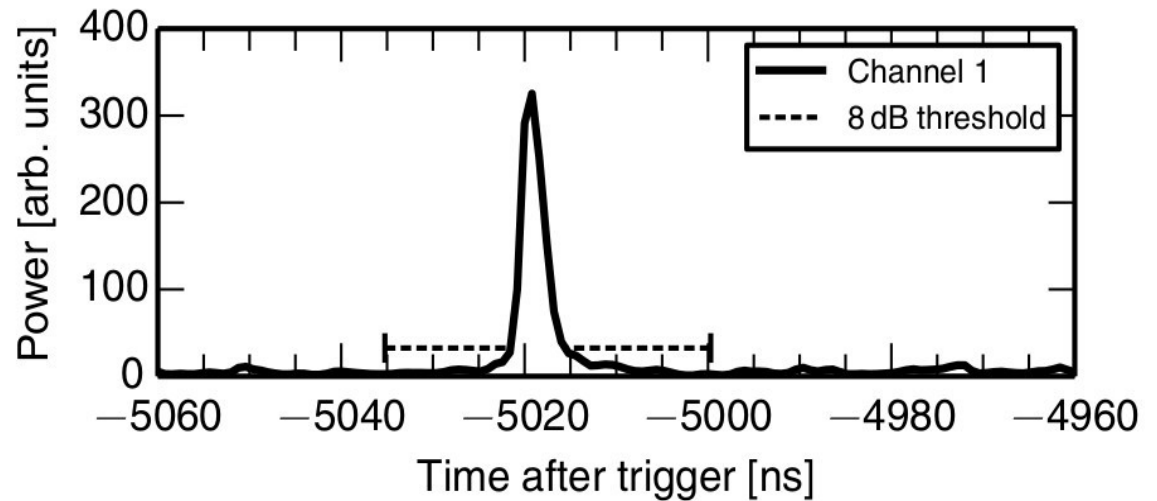
An example of the recorded signal (shower energy  $9 \times 10^{16}\ \text{eV}$ , shower axis 126 m from the antenna).



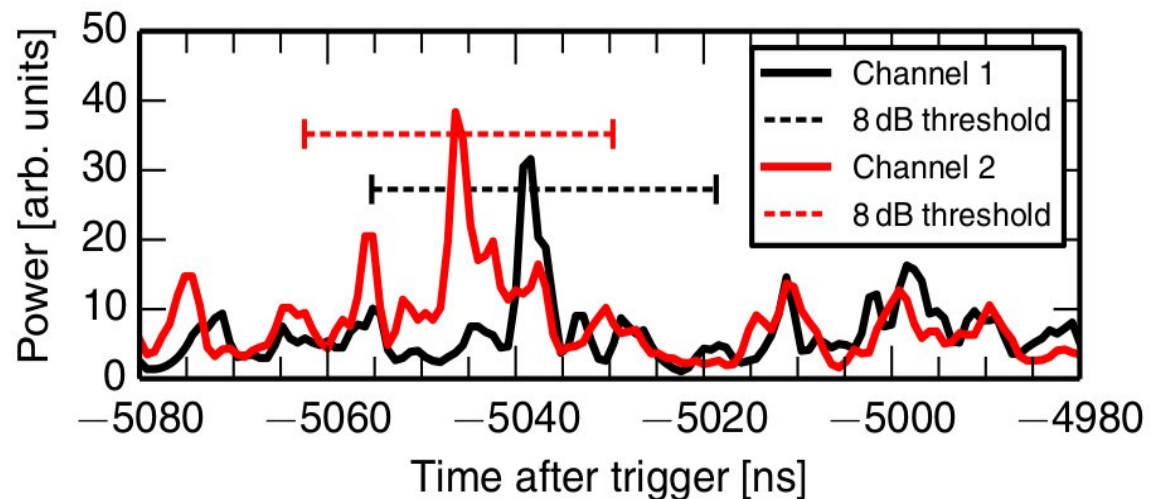
# Results – examples

Event selection – signal at least 8 dB above the mean noise level.

Highest signal: 17.7 dB  
Energy =  $2.5 \times 10^{17}$  eV  
**Zenith angle =  $5.6^\circ$**   
**Distance = 120 m**



Stereo event  
Signals: 8.2 and 8.5 dB  
Energy =  $3.7 \times 10^{16}$  eV  
**Zenith angle =  $3.7^\circ$**   
**Distance = 110 m**

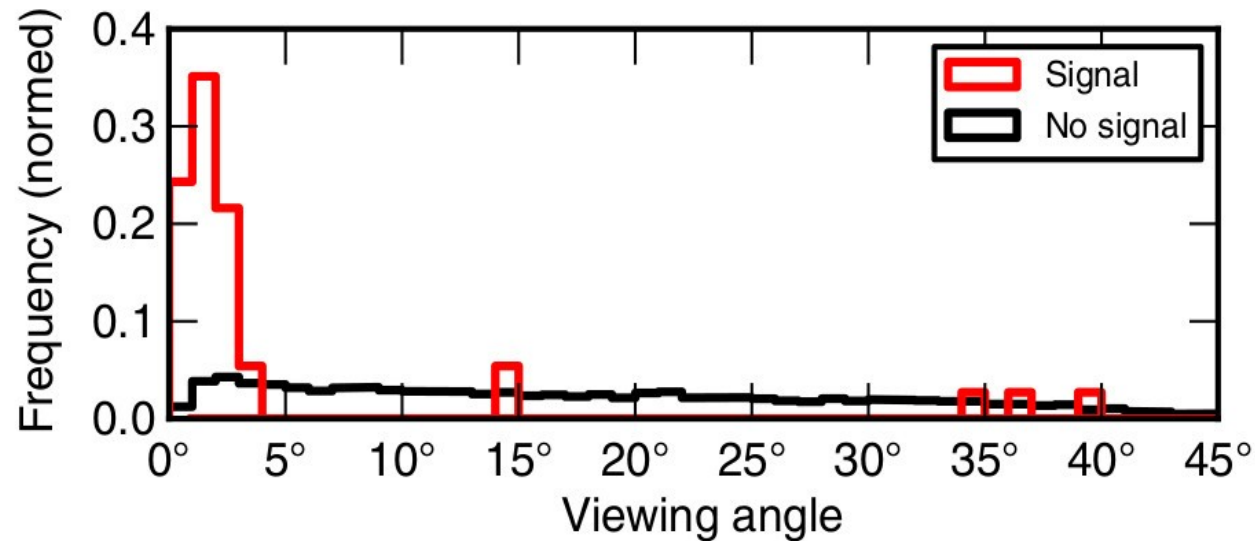




# Results – event selection

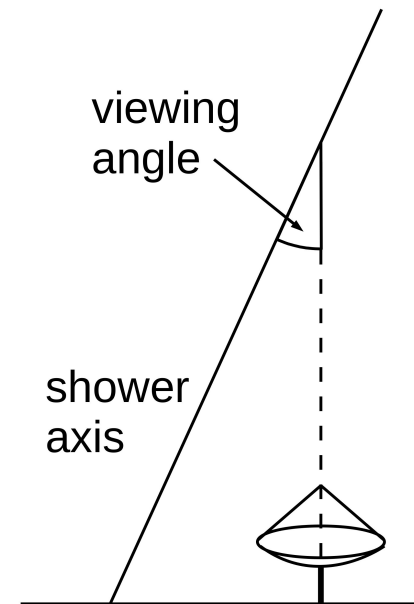
Cuts on data used in analysis:

- Shower energy  $> 3 \times 10^{16}$  eV,
- Shower axis within field of view of at least one receiver ( $2^\circ$ ),
- Signal within  $\Delta t \approx 50$  ns from expected arrival time,
- Signal  $> 8$  dB above noise.



With these conditions applied, there are 37 event candidates with microwave signal from the KASCADE-Grande showers. Expected number of noise signals is  $9.4 \pm 0.2$ .

After introducing a cut on viewing angle  $< 4^\circ$ : 31 showers with microwave signal. Expected number of noise events is only  $1.1 \pm 0.1$ .



# Radio emission from extensive air showers

Sources of emission:

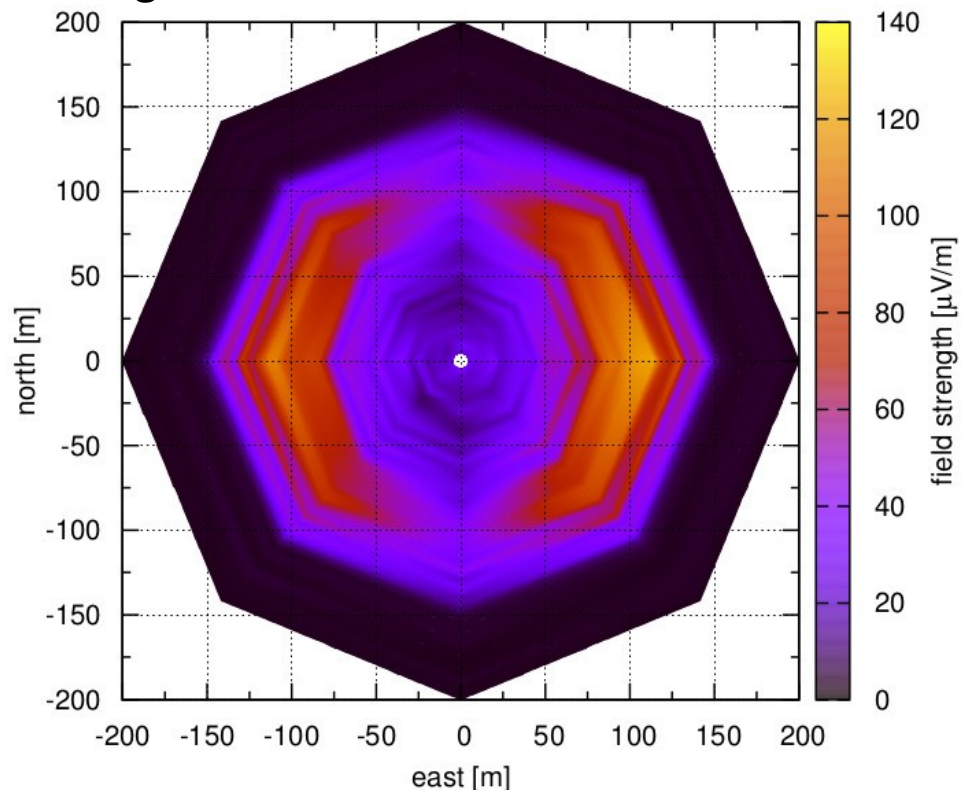
- charged particles deflected in Earth's magnetic field – geomagnetic radiation,
- more electrons than positrons in air shower – Askaryan effect.

**Signal observed in the MHz range, but suppressed for shorter wavelengths.**

**Near the Cherenkov angle, time compression effects are expected – this should enable observations in the GHz range.**

CoREAS – CORSIKA-based program for simulation of radio signals from extensive air showers.

Predicted signal in the 3.4-4.2 GHz band, from a vertical  $10^{17}$  eV shower induced by a proton.





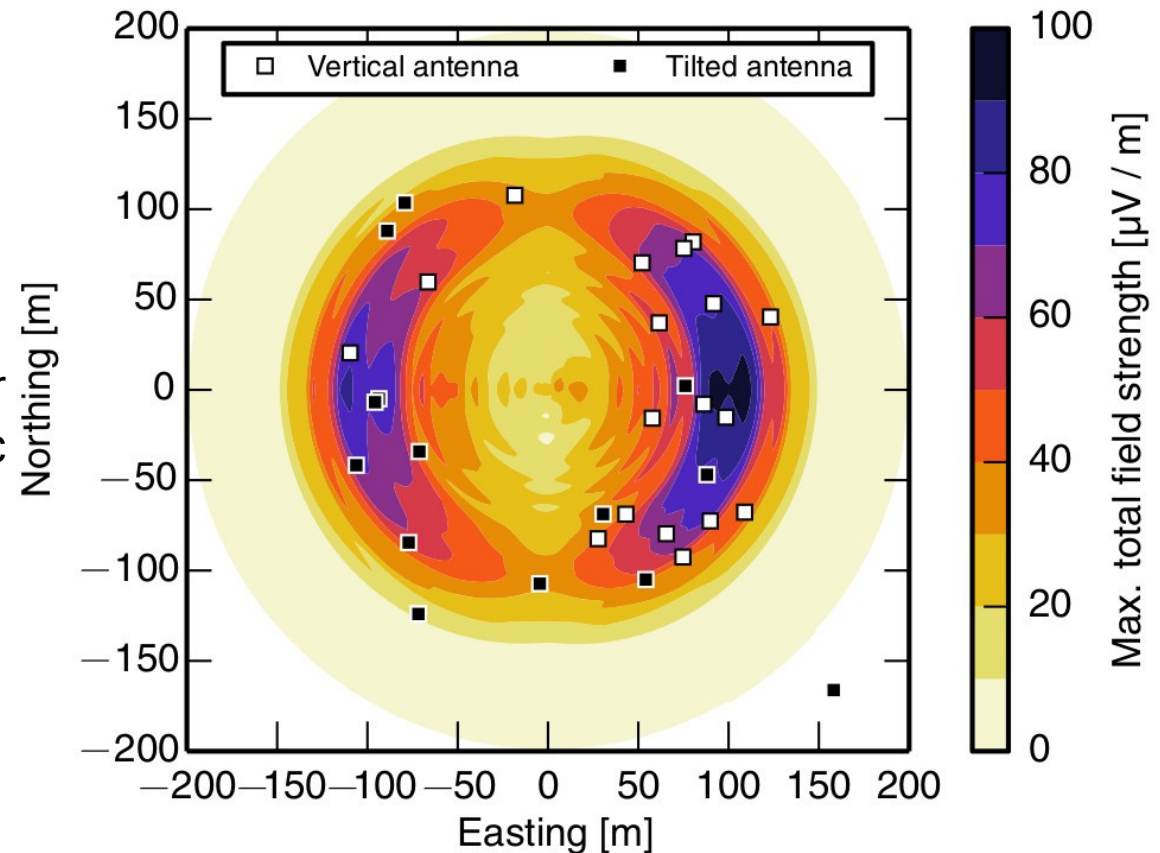
# Results – interpretation

Reconstructed positions of shower cores form a ring structure around the antennas.

High signals at the Cherenkov angle relatively to the air shower axis – **consistent with time compressed geomagnetic radio emission.**

Stronger signals for showers with larger angles relative to the local geomagnetic field.

East-west asymmetry in the field strength – addition from Askaryan effect?



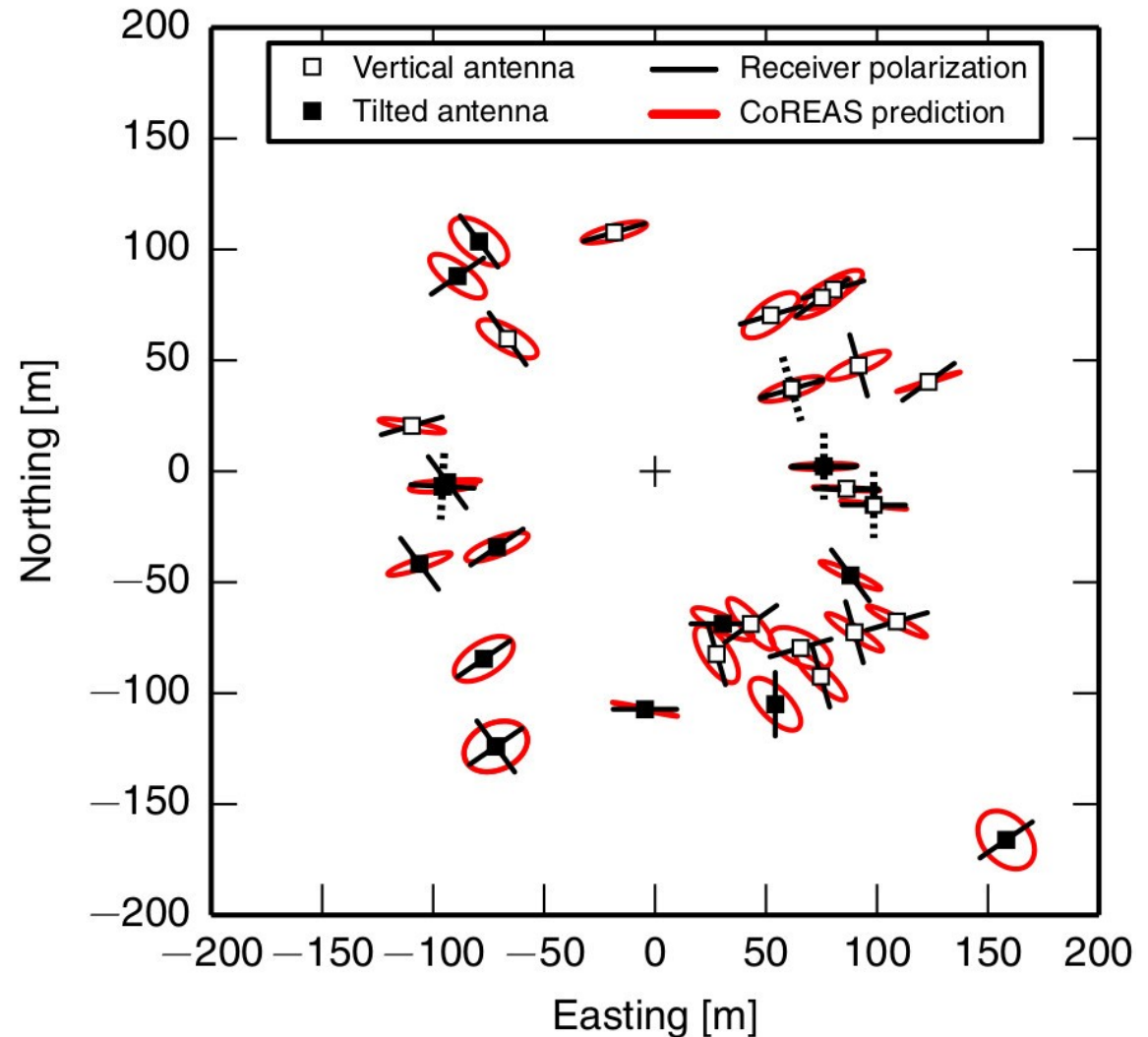
CoREAS prediction for a vertical shower  $10^{17}$  eV and  $X_{\max} = 658$  g/cm<sup>2</sup>

# Results – polarization

Polarisation pattern consistent with the geomagnetic radio emission.

Statistical analysis rejects the unpolarized microwave emission with  $5\sigma$  significance.

(Molecular bremsstrahlung is expected to be unpolarized.)



Polarization directions of receivers with microwave signal and the CoREAS prediction.



# Summary

CROME recorded microwave emission from over 30 extensive air showers.

The measurements show that microwave radiation offers a new means of studying air showers.

Evidences for time compressed geomagnetic and Askaryan radio emission:

- small viewing angle (close to the Cherenkov cone) relative to air shower axis,
- dependence on the angle relative to geomagnetic field,
- east-west asymmetry.

Measured microwave signal is polarized.

**⇒ The molecular bremsstrahlung is NOT the dominant emission mechanism of the microwave signals.**

More details in Phys. Rev. Lett. 113, 221101 (2014)